# NIST DTSA-II ("Son of DTSA"): Step-by-Step

### Dale E. Newbury (grateful user) National Institute of Standards and Technology Gaithersburg, MD 20899-8370

### **NIST-NIH Desktop Spectrum Analyzer (DTSA)**

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### **Spectral Simulation with DTSA**



Photon Energy (keV)

For 16 years, I have heard: "When will you have DTSA for the pc?"

 Created by Nicholas Ritchie of NIST (nicholas.ritchie@nist.gov), inspired by NIST-NIH Desktop Spectrum Analyzer (DTSA) invented 1990-92 by Chuck Fiori (NIH and NIST) and Carol Swyt-Thomas (NIH), and then further developed by Carol and Bob Myklebust (NIST).

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- DTSA-II is being continually improved and the latest version can be downloaded for free at http://www.cstl.nist.gov/div837/837.02/epq/dtsa2/index.html

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Chemical Science and Te Surface and Micr	chnology Laboratory roanalysis Science Division	National Institute of Standards and Technology						
DTSA-11			Son of DeskTop Spectrum	Analyzer				
Geek								
DTSA-II is a multiplatform software package for quantitative x-ray microanalysis. DTSA-II was inspired by the popular Desktop Spectrum Analyzer (DTSA) package developed by Chuck Fiori, Carol Swyt-Thomas, and Bob Myklebust at NIST and NIH in the '80's and early '90's.								
Introduction Installation	DTSA-II has being designed with the goal of ma microanalyst. <i>We want to encourage standard</i> . Many operations which had previously required	iking standards-based microanalysis more s-based analysis by making it as easy a user intervention under DTSA now are pe	accessible for the novice s possible to get reliable results erformed entirely by the software.					
<u>Getting Started</u> <u>Simulation</u> <u>Quantification</u>	sanity checks. While this might not provide the fluis more consistent with the way laboratories are experts in single techiques. We encourage users	exibility that some sophisticated users may moving towards technicians responsible for who desire the additional power and flexi	sses while performing quarty con y desire, we feel that this philosop or multiple techniques and away fi bility available in the EPQ library	hy irom to				
<u>Why Java?</u> <u>3rd Party Licenses</u>	learn to script using Jython or to create their own exposed via DTSA-II.	n alternative user interface. EPQ is much r	nore capable than the fraction					
<u>Contact Us</u>	DTSA-II is based on an entirely new code base into a shared algorithm library which forms the b the user interface shell and the EPQ library is the	written by Nicholas W. M. Ritchie. The o asis for a handful of software products an algorithm library.	odebase has been carefully divid d a user interface shell. DTSA-II	ed is				
NIST Home Page	CSTL Home Page	SMSD Home Page	NISTMonte Home Pag	ge				
DISCLAIMER: Any mention	n of commercial products is for information on	ly; it does not imply recommendation o	r endorsement by NIST nor doe	es it imply tha				
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- DTSA-II is being continually improved and the latest version can be downloaded for free at http://www.cstl.nist.gov/div837/837.02/epq/dtsa2/index.html
- Tools currently embedded in DTSA-II:

### MAJOR FEATURES:

#### **Basic IO and Display**

- Read energy dispersive x-ray spectra in a variety of different commercial and non-commercial formats including the industry standard EMSA format
- o Display and overlay spectra with various scaling options on linear/log/sqrt axes
- o Copy/save/print the spectrum display as a bitmap/PNG file
- o Output the spectra as a GNUPlot file for publication quality output
- o Overlay labeled x-ray emission lines and x-ray absorption edges
- Define and integrate regions-of-interest
- View spectrum contextual information
- o Archive spectra to a searchable database
- o Sub-sampling of spectral data to simulate shorter acquisition times

### **Basic operations**

• Opening and manipulating spectral files



#### Display and overlay spectra with various scaling options on linear/log/sqrt axes



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#### Display and overlay spectra with various scaling options on linear/log/sqrt axes



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### But wait!



# But wait!

• A brilliant feature, the "Report" is going to record your actions. A daily diary of actions (file named by date) is automatically saved.

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## **Basic operations**

- Opening and manipulating spectral files
- Display of spectra





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### **Basic operations**

- Opening and manipulating spectral files
- Display of spectra
- Peak labeling (manual only)











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### **Basic operations**

- Opening and manipulating spectral files
- Peak labeling (manual only)
- Exporting spectra for publication (Gnuplot)

#### **Exporting spectrum for publication as gnuplot**



#### Publication-quality graphics from Gnuplot



### Comparing spectra from different EDS spectrometers, or from different dates from the same EDS: The issue of EDS calibration

- A spectrum is recorded with calibration data: eV/channel, zero offset, number of channels (depending on manufacturer, this data may or may not be embedded in the .msa header).
- When a spectrum is read into DTSA-II, the calibration information is checked against the current calibration. If there is a mismatch, a message prompts the analyst.

#### **Spectrum Calibration**



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## The issue of EDS calibration

- A spectrum is recorded with calibration data: eV/channel, zero offset, number of channels (depending on manufacturer, this data may or may not be embedded in the .msa header).
- When a spectrum is read into DTSA-II, the calibration information is checked against the current calibration. If there is a mismatch, a message prompts the analyst.
- The analyst then has two choices:
  - 1. Change the detector selection to match the incoming spectrum
  - 2. Accept the incoming spectrum but display it according to the current calibration information. (Note: the incoming spectrum will retain its calibration data so that when it is the only spectrum being displayed, its own calibration will be applied.)

• Matching to a particular ROI





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### **Spectrum sub-sampling tool**

- Take an experimentally measured spectrum and create one or more "sub-samples", that is, equivalent spectra that would have been collected at lower dose.
- Sub-sampled spectra are useful for statistical studies. e.g., how does detection limit vary with dose.









## Background fitting tool













# **DTSA-II Simulation Mode**

- EDS spectra calculated from
  - 1. First principles, using best available cross sections and physical data (flat, bulk target only)

#### **Simulation Alien**

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#### Simulation Alien: target composition



#### **Simulation Alien: instrument configuration**



#### **Simulation Alien: Other options not invoked**


## **Simulation Alien**



### **Simulation Alien**



## **Simulation Alien: Other options invoked**



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## **Simulation Alien**

## Simulation Alien



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# **DTSA-II Simulation Mode**

- EDS spectra calculated from
  - 1. First principles, using best available cross sections and physical data (flat, bulk target only)
  - 2. Monte Carlo electron trajectory simulation for various specimen configurations:
    - 1. Flat, bulk
    - 2. Layer on bulk
    - 3. Inclusion (hemisphere) embedded in bulk
    - 4. Spherical particle on substrate
    - 5. Cubic particle on substrate



Detector thickness = 5 mm



Detector thickness = 5 mm

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## Simulation Alien: Monte Carlo simulation trajectories can be viewed with Cosmo Player

	Player 2.1.1       Getting Started       Quick Reference       Going Further       Release Notes       About Cosmo Player						
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<u>Player</u> <u>2.1.1</u>							
<u>On the</u> Dashboard	Cosmo Player plugs in to your Web browser to enable you to see and explore 3D worlds. With Cosmo Player you can visit any 3D world authored in the Virtual Reality Modeling Language (VRML). These 3D worlds often include other kinds of multimedia, like sound and movies.						
<u>Moving</u> <u>Around in</u> <u>a World</u>	This brief guide shows you the basics of the Cosmo Player main controls so you can get started right away.						
Examining Objects	You can find more in-depth information in Cosmo Player 2.1.1 Quick Reference and Going Further with Cosmo Player 2.1.1. You can find more technical information about installation and trouble-shooting in the Release Notes.						
<u>Changed</u> <u>Your</u> <u>Mind?</u>	You can practice using Cosmo Player by playing CHOMPY, an interactive 3D game that teaches basic navigation.						
Interacting with Active Objects							
Another	On the Dashboard						
Way of Moving Through a World	You use the main controls on the Cosmo Player dashboard to do two things: move around in 3D worlds and examine objects in 3D worlds. (Some worlds don't display the dashboard, but they may provide on-screen cues to navigation.)						





## View along the beam



## View along the beam



## View from bottom of particle



# **DTSA-II Simulation Mode**

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  - 1. First principles, using best available cross sections and physical data (flat, bulk target only)
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    - 5. Cubic particle on substrate



Detector thickness = 5 mm



Detector thickness = 5 mm





## **DTSA-II:** Quantitative Analysis

- ZAF analysis against standards
- Standards are used to extract needed peak references for MLLS fit.
- Report contains pertinent data (ZAF factors, weight%, atom%, normalized weight%; 1σ statistics)

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Default Detector     SF     Oxygen by stoichiometry     Element     Cation     Anion     As       8600 Probe     Image: Second State	nes Temporary Lines Fe Selected Lines
Spectrum List	Lines Clear All Lines
Message:     Specify an element to be handled specially     More       Back     Next     Finish     C	ancel
None         All         Clear           Welcome to NIST DTSA-II version 1698	



NIST DTSA-II			_ <b>_</b> X
Elle Process Tools Help			
15-			
<b>1</b>			
Quantification Alien			×
→I <b>←</b>	Previor	s: Select the unknown spectrum	
The r	esults		
T.		Finish	
**			
5 10 Normalization			
Spectrum Report   Command	C Normalized weight percent	Atomic percent	
Default Detector Spectrum	Sum	5 Fe	[ ]
8600 Probe	. 100.10 ± 0.26  36.41 ± 0.0	9 03.09 ± 0.25	ines <u>E</u> lement: Temporary Lines
EDAX_35mus			nes Fe Selected Lines
Spectrum List			Lines Clear All Lines
			sition
Message:		More	ment Weight % Atomic %
	Back	ext Finish Cance	
None All Clear			
Welcome to NIST DTSA-II version 1698			
MIST DTSA-II File Process Tools Help			
---	---	---	
1 🔶 10-	Quantification Alien	X	
→I <b>←</b> ↓ 5-	Previous: Select the unknown spectrum		
++	Finish		
5 10 15 20 0 2,0	Normalization           O         Weight percent         C         Normalized weight percent         E         Atomic percent	16,000 18,000 20,0	
Spectrum Report Command Default Detector 8600 Probe	Spectrum         Sum         S         Fe           Fe5-20kV750pA35mu         100.10 ± 0.26         49.89 ± 0.22         50.11 ± 0.26	nes ines <u>E</u> lement: Temporary Lines	
EDAX_35mus		Ines Fe Selected Lines	
r.	Message: More Back Mext Finish Cancel	sition 	
r.			
None All Clear			











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NIST DTSA-II			
Eile Process Tools Help			
Cu-stdB_20kV20	JnAMed298kHz21DT100s		
20,000- User Information			
Quantitative algorithms     Detector - BrukerQuadSDD_145eV     Detector - QUAD SDD 5eV     Detector - QUAD SDD 5eV			
Detector - BrukerQuadSDD_10eV     Detector - BrukerQuadSDD_128eV     Detector - BrukerQuadSDD_145eV     Azimuthal angle 180.0 °			
Image: Second			
5,000- 5,000- Crystal parameters			
Detector Area 40.0 mm <sup>2</sup>	9,000 10,0		
Spectrum     Report     Comm       Default Detector     Aluminum layer     10.0	1		
JEOL8500F         Dead layer         0.0         µm           BrukerQuadSDD-10eV         Thickness         0.4         mm	Temporary Lines		
Spectrum List Cu-stdB_20kV20nAMed29 Edit the properties of this detector.	Clear All Lines		
OK Cancel Apply ig	ht % Atomic %		
Detector orientation     [0.893,-0.000,0.451]       Detector position     [-61.436,0.000,-31.018]       Detector thickness     0.45 mm       Detector type     Silicon Drift Detector			
None     All     Clear     Detector window     Moxtek AP 3.3 (manufacturer's table)       Welcome to NIST DTSA-II version 1705			

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Eile Process Iools Help				
25,000-	Cu-stdB_20kV20nAMed298kHz21DT100s			
20,000-				
User Information User Information Quantitative algorithms Instruments and Detectors I5,000- ISOURSOUF Detector - QUAD SDD SeV	Detector - BrukerQuadSDD_145eV			
Detector - BrukerQuadSDD-10eV     Detector - BrukerQuadSDD_128e'	Aluminum layer 10.0 nm			
10,000- 10,	Thickness 0.4 mm			
	Number of channels 4096 channels			
	Zero strobe discriminator 50.0 eV			
0 Spectrum Report Comm	Base Performance 9,000 10,0			
Default Detector	Zero offset -475.0 eV			
BrukerQuadSDD-10eV	Resolution 145.0 eV at Mn Ka			
Spectrum List Cu-stdB_20kV20nAMed29	Edit the properties of this detector.			
	OK Cancel Apply ight % Atomic %			
Detector orientation Detector position	[0.893,-0.000,0.451] [-61.436,0.000,-31.018]			
Detector thickness	U.45 MM Silicon Drift Detector			
None All Clear Detector window	Moxtek AP 3.3 (manufacturer's table)			
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